Lecture 2

The Semi Empirical Mass Formula SEMF

Overview

1 The liquid drop model 2 The Coulomb Term 3 Mirror nuclei, charge asymmetry and independence 4 The Volume and Surface Terms 5 The asymmetry term 6 The pairing term 7 The SEMF

0 Introduction to the SEMF

- Aim: phenomenological understanding of nuclear binding energies as function of A, Z and N.
 Assumptions:
 - Nuclear density is constant (see lecture 1).
 - We can model effect of short range attraction due to strong interaction by a liquid drop model.
 - Coulomb corrections can be computed using electro magnetism (even at these small scales)
 - Nucleons are fermions at T=0 in separate wells (Fermi gas model → asymmetry term)
 - QM holds at these small scales → pairing term.
- Compare with experiment: success & failure!

1 Liquid Drop Model Nucleus

- Phenomenological model to understand binding energies.Consider a liquid drop
 - Ignore gravity and assume no rotation
 - Intermolecular force repulsive at short distances, attractive at intermediate distances and negligible at large distances → constant density.
 - *n*=number of molecules, *T*=surface tension, *B*=binding energy *E*=total energy of the drop, α,β =free constants

 $E = -\alpha n + 4\pi R^2 T \Rightarrow B = \alpha n - \beta n^{2/3}$ surface area ~ n^{2/3}

- Analogy with nucleus
 - Nucleus has constant density
 - From nucleon-nucleon scattering experiments we know:
 - Nuclear force has short range repulsion and is attractive at intermediate distances.
 - Assume charge independence of nuclear force, neutrons and protons have same strong interactions → check with experiment (Mirror Nuclei!)

2 Coulomb Term

- The nucleus is electrically charged with total charge Ze
- Assume that the charge distribution is spherical and compute the reduction in binding energy due to the Coulomb interaction

$$E_{Coulomb} = \int_{0}^{2e} \frac{Q(r)}{4\pi\varepsilon_0 r} dQ \qquad Q(r) = Ze(r/R)^3 \quad dQ = 3Zer^2/R^3 dr$$

to change the integral to dr; R=outer radius of nucleus

 $E_{Coulomb} = \int_{0}^{R} \frac{3(Ze)^{2}}{4\pi\varepsilon_{0}r} \frac{r^{5}}{R^{6}} dr = (3/5) \frac{(Ze)^{2}}{4\pi\varepsilon_{0}R}$... and remember $R = R_{0}A^{-1/3}$ in calculation of the second state of the second s

includes self interaction of last proton with itself. To correct this replace Z^2 with $Z^*(Z-1)$

in principle you could take d from this calculation but it is more accurate to take it from the overall fit of the SEMF to data (nuclei not totally spherical or homogeneous)

3 Mirror Nuclei

- Does the assumption of the drop model of constant binding energy for every constituent of the drop acatually hold for nuclei?
- Compare binding energies of mirror nuclei (nuclei with n←→p). Eg ⁷₃Li and ⁷₄Be.
- If the assumption holds the mass difference should be due to n/p mass difference and Coulomb energy alone.
- Let's compute the Coulomb energy correction from results on previous page

$$\Delta E_{coulomb}(Z, Z-1) = \frac{3}{5} \frac{e^2}{4\pi\varepsilon_0 R} [Z(Z-1) - (Z-1)(Z-2)] = \frac{3}{5} \frac{e^2}{4\pi\varepsilon_0 R} 2(Z-1)$$

 $Z \sim A/2$; $R = R_0 A^{1/3}$ to find that $\Delta E_C(Z, Z-1) \propto A^{2/3}$

- Now lets measure mirror nuclei masse, assume that the model holds and derive $\Delta E_{Coulomb}$ from the measurement.
- This should show an A^{2/3} dependence
- And the scaling factor should yield the correct R_0 of 1.2 fm
- if the assumptions were right

Charge symmetry

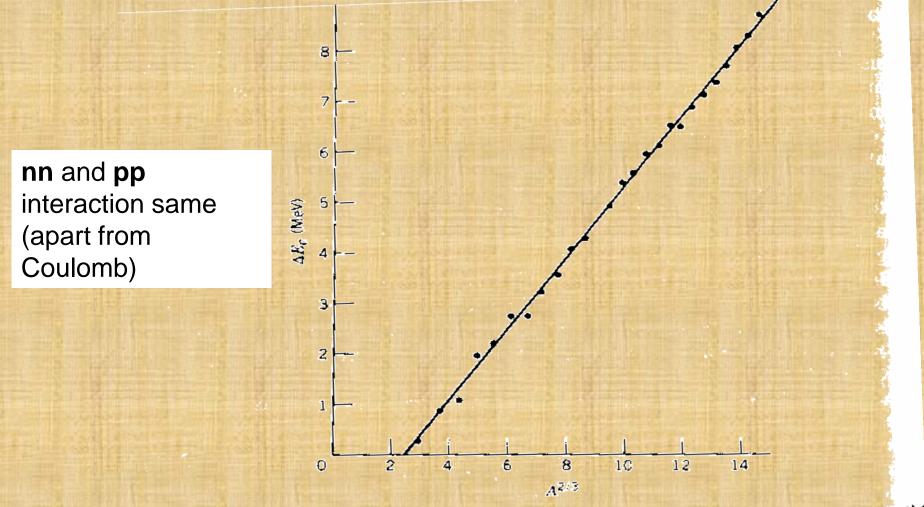
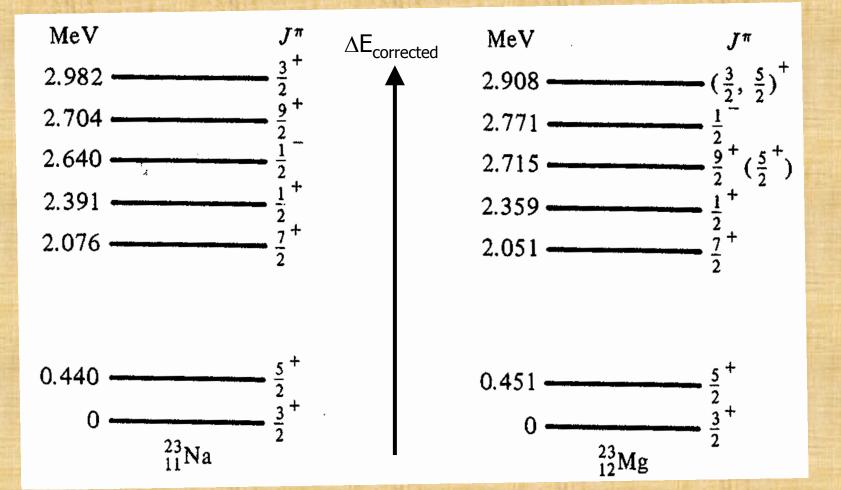


Figure 3.10 Coulomb energy differences of mirror nuclei. The data show the expected $A^{2/3}$ dependence, and the slope of the line gives $R_0 = 1.22$ fm.

3 More charge symmetry

Energy Levels of two mirror nuclei for a number of excited states
Corrected for n/p mass difference and Coulomb Energy



3 From Charge Symmetry to Charge Independence

- Mirror nuclei showed that strong interaction is the same for nn and pp.
- What about np ?
- Compare energy levels in "triplets" with same A, different number of n and p. e.g.

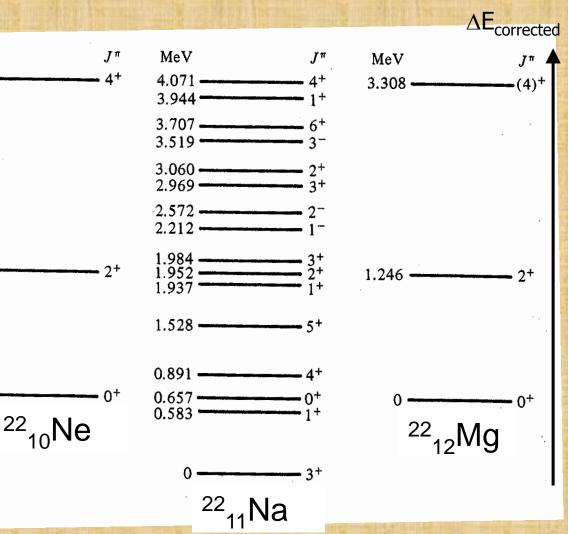


If we find the same energy levels for the same spin states
 Strong interaction is the same for np as nn and pp.
 See next slide

3 Charge Independence

MeV

- Same spin/parity states should have the same 3.357 energy.
- Yes: np=nn=pp
- Note: Far more states in ²²₁₁Na. Why?
- Because it has more np pairs then the others
- np pairs can be in any 1.275 • Spin-Space configuration
- pp or nn pairs are excluded from the totally symmetric ones by Herr Pauli
- Note also that ²²₁₁Na has the lowest (most bound) state, remember for the deuteron on next page



3 Charge Independence

We have shown by measurement that:

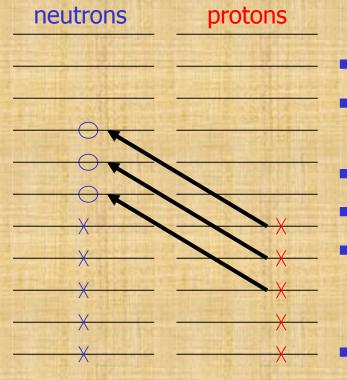
- If we correct for n/p mass difference and Coulomb interaction, then energy levels in nuclei are unchanged under n ← → p
- and we must change nothing else! I.e. spin and space wavefunctions must remain the same!
- Conclusion: strong two-body interaction same for pp, pn and nn if nucleons are in the same quantum state.
- Beware of the Pauli exclusion principle! eg why do we have bound state of pn but not pp or nn?
 - because the strong force is spin dependent and the most strongly bound spin-space configurations (deuteron) are not available to nn or pp. It's Herr Pauli again!
 - Just like ²²₁₁Na on the previous triplet level schema

4 Volume and Surface Term

- We now have all we need to trust that we can apply the liquid drop model to a nucleus
 - constant density
 - same binding energy for all constituents
- Volume term: $B_{Volume}(A) = +aA$ Surface term: $B_{Surface}(A) = -bA^{2/3}$
- Since we are building a phenomenological model in which the coefficients a and b will be determined by a fit to measured nuclear binding energies we must inlcude any further terms we may find with the same A dependence together with the above

 ■ Neutrons and protons are spin ½ fermions → obey Pauli exclusion principle.

 If all other factors were equal nuclear ground state would have equal numbers of n & p.



Illustration

- n and p states with same spacing Δ .
- Crosses represent initially occupied states in ground state.
- If three protons were turned into neutrons
- the extra energy required would be $3 \times 3 \Delta$.
- In general if there are Z-N excess protons over neutrons the extra energy is $((Z-N)/2)2 \Delta$. relative to Z=N.

But how big is Δ ?

Assume:

- p and n form two independent, non-interacting gases occupying their own square Fermi wells
 kT << Δ
- so we can neglect kT and assume T=0
- This ought to be obvious as nuclei don't suddenly change state on a warm summers day!
- Nucleons move non-relativistically (check later if this makes sense)

From stat. mech. density of states in 6d phase space = 1/h³

 $dN_{particle} = \frac{4\pi p^2 dpV}{h^3}$ here N_{particle} could be the number of protons or neutrons

 Integrate up to p_f to get total number of protons Z (or Neutrons N), & Fermi Energy (all states filled up to this energy level).

$$Z = (4\pi V / 3h^3) p_F^3$$
 and $V = \frac{4}{3}\pi R_0^3 A$

$$P_F = (3/4\pi)^{2/3} \frac{h}{R_0} \left(\frac{Z}{A}\right)^{1/3} \text{ and } E = \frac{p^2}{2m} \Rightarrow E_F = (3/4\pi)^{4/3} \frac{h^2}{2mR_0^2} \left(\frac{Z}{A}\right)^{2/3}$$

• Change variables $p \rightarrow E$ to find avg. E $dN / dE = \frac{dN / dp}{dp / dE} = const \cdot E^{1/2} \quad \langle E \rangle = \frac{0}{\frac{0}{E_F}} E^{3/2} dE = (3/5)E_F$

These are all standard stat. mech. results!

• Compute total energy of all protons by $Z^* < E >$ call this K

$$E_{Total}^{P} = Z \left\langle E^{P} \right\rangle = \frac{3}{5} (3/4\pi)^{4/3} \frac{h^{2}}{2mR_{0}^{2}} Z \left(\frac{Z}{A}\right)^{2/3}$$

Use the above to compute total energy of Z protons and N neutrons

 $E_{Total} = \frac{K}{A^{2/3}} \begin{bmatrix} Z^{5/3} + N^{5/3} \end{bmatrix}$ change variables from (Z,N,A) to (y,A) with y=N-Z

 $E_{Total} = \frac{KA^{5/3}}{A^{2/3} 2^{5/3}} \Big[(1 - y/A)^{5/3} + (1 + y/A)^{5/3} \Big] \text{ where y/A is a small number (ϵ)}$ $E_{Total} = \frac{KA}{2^{5/3}} \Big[(1 - \varepsilon)^{5/3} + (1 + \varepsilon)^{5/3} \Big] (1 - \varepsilon)^{5/3} = 1 - \frac{5}{2} \varepsilon + \frac{52}{2} \varepsilon^2 + O(\varepsilon^3)$

 $2^{5/3} \lfloor (1-\epsilon)^{5/3} \rfloor = 1 - \frac{5}{3}\epsilon + \frac{52}{33}\epsilon^2 + O(\epsilon^3)$ Binomial expansion keep lowest term in y/A $E_{Total} = 2^{-2/3} KA + \frac{2^{1/3}5}{9} K \frac{(N-Z)^2}{A} \Delta E_{Total} (Fermi-Gas) = \frac{const}{4} * \frac{(N-Z)^2}{A}$

This terms is only proportional to volume (A). It has already been captured by the Volume term of the liquid drop model

comes from a fit of the SEMF to measurements analytical ≈ 24 MeV

From the Fermi Gas model we learn that

 due to the fermionic nature of p and n we loose in binding energy if the nucleus deviates from N=Z

• The Asymmetry term: $B_{Asymmetry}(N,Z) = -c \frac{(N-Z)^2}{A}$

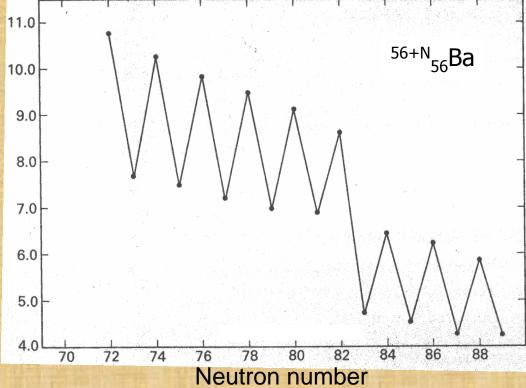
Observations:

- Nuclei with even number of n or even number of p more tightly bound then with odd numbers. See figure
- Only 4 stable o-o nuclei but 153 stable e-e nuclei.
- p energy levels are Coulomb shifted wrt n → small overlap of wave functions between n and p.
- Two p or two n in same energy level with opposite values of j_z have AS spin state
 - forced into sym spatial w.f.
 maximum overlap
 - maximum binding energy because of short range attraction.

Note: this only holds for nn and pp, not for np. → We don't have a preference for even A

6 Pairing Term

Neutron separation energy [MeV] in Ba isotopes



6 Pairing Term

Measure that the Pairing effect smaller for larger A
 Phenomenological*) fit to A dependence gives A^{-1/2}

R $(A) -$	<u> </u>
$B_{Pairing}(A) =$	$\overline{A^{1/2}}$

	δ
e-e	+ive
e-o	0
0-0	-ive

Note: If you want to plot binding energies versus A it is often best to use odd A only as for these the pairing term does not appear

> *) For an even more insightful explanation of the A dependence read the book by Jelley

6 Semi Empirical Mass Formula

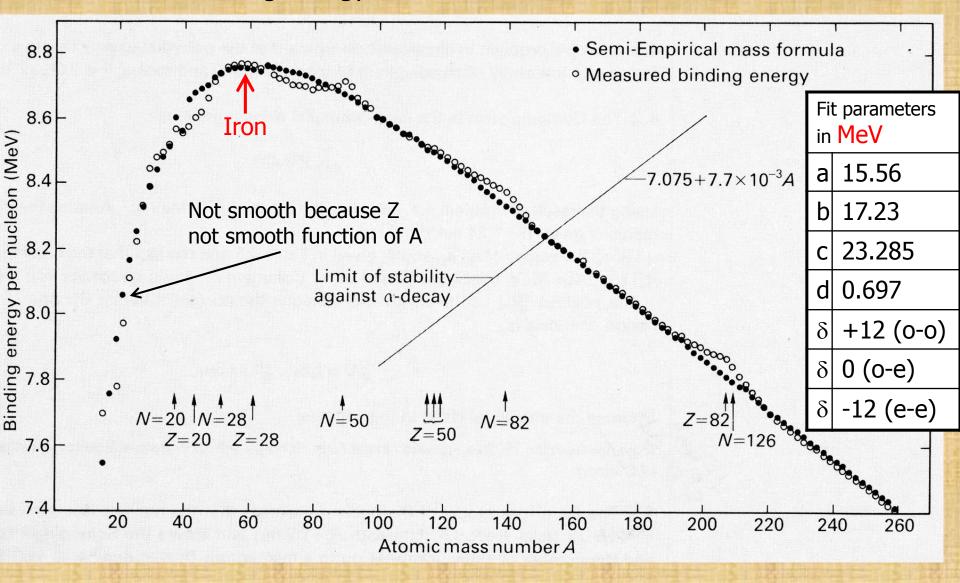
Put everything together:

Lets see how all of these assumptions fit reality

- And find out what the constants are
 - Note: we went back to the simpler Z² instead of Z*(Z-1)

6 Semi Empirical Mass Formula

Binding Energy vs. A for beta-stable odd-A nuclei



6 Semi Empirical Mass Formula

Conclusions

- Only makes sense for A≥20
- Good fit for large A (good to <1%) in most places.
- Deviations are interesting → shell effects.
- Coulomb term constant agrees with calculation.
- Explains the valley of stability (see next lecture).
- Explains energetics of radioactive decays, fission and fusion.